

Solving DSGE Models

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Solving DSGE Models

- A variety of numerical techniques are available to solve rational expectations systems.
- In the context of likelihood-based DSGE model estimation linear approximation methods are very popular because they lead to a state-space representation of the DSGE model that can be analyzed with the Kalman filter.
- Log-linearization of $f(x)$:
 1. write $f(x) = f(e^z)$;
 2. conduct a first-order Taylor approximation around x_0 in terms of z :

$$f(e^{\ln x}) \approx f(x_0) + x_0 f^{(1)}(x_0)(\ln x - \ln x_0).$$

Solving DSGE Models

- Our DSGE model leads to:

$$\hat{y}_t = \mathbf{IE}_t[\hat{y}_{t+1}] + \hat{g}_t - \mathbf{IE}_t[\hat{g}_{t+1}] - \frac{1}{\tau} \left(\hat{R}_t - \mathbf{IE}_t[\hat{\pi}_{t+1}] - \mathbf{IE}[\hat{z}_{t+1}] \right) \quad (1)$$

$$\hat{\pi}_t = \beta \mathbf{IE}_t[\hat{\pi}_{t+1}] + \kappa(\hat{y}_t - \hat{g}_t) \quad (2)$$

$$\hat{c}_t = \hat{y}_t - \hat{g}_t, \quad (3)$$

$$\hat{R}_t = \rho_R \hat{R}_{t-1} + (1 - \rho_R) \psi_1 \hat{\pi}_t + (1 - \rho_R) \psi_2 (\hat{y}_t - \hat{g}_t) + \epsilon_{R,t} \quad (4)$$

$$\hat{g}_t = \rho_g \hat{g}_{t-1} + \epsilon_{g,t} \quad (5)$$

$$\hat{z}_t = \rho_z \hat{z}_{t-1} + \epsilon_{z,t} \quad (6)$$

where

$$\kappa = \tau \frac{1 - \nu}{\nu \pi^2 \phi}.$$

- Introduce $\eta_t^\pi = \hat{\pi}_t - \mathbf{IE}_{t-1}[\hat{\pi}_t]$ and $\eta_t^y = \hat{y}_t - \mathbf{IE}_{t-1}[\hat{y}_t]$.

Solving DSGE Models

- Linearized DSGE leads to linear rational expectations (LRE) system:

$$\Gamma_0(\theta)s_t = \Gamma_1(\theta)s_{t-1} + \Psi\epsilon_t + \Pi\eta_t \quad (7)$$

where

- s_t is a vector of model variables, ϵ_t is a vector of exogenous shocks,
 - η_t is a vector of RE errors with elements $\eta_t^x = \hat{x}_t - \mathbf{IE}_{t-1}[\hat{x}_t]$, and
 - s_t contains (among others) the conditional expectation terms $\mathbf{IE}_t[\tilde{x}_{t+1}]$.
- Solution methods for LREs: Blanchard and Kahn (1980), King and Watson (1998), Uhlig (1999), Anderson (2000), Klein (2000), Christiano (2002), Sims (2002).
 - Overall the solution in terms of s_t is of the form

$$s_t = \Phi_1(\theta)s_{t-1} + \Phi_\epsilon(\theta)\epsilon_t. \quad (8)$$

Example

- Consider the simplified system:

$$\hat{y}_t = \mathbf{IE}_t[\hat{y}_{t+1}] - \frac{1}{\tau}(\hat{R}_t - \mathbf{IE}_t[\hat{\pi}_{t+1}]) \quad (9)$$

$$\hat{\pi}_t = \beta \mathbf{IE}_t[\hat{\pi}_{t+1}] + \kappa \hat{y}_t \quad (10)$$

$$\hat{R}_t = \psi_1 \hat{\pi}_t + \epsilon_{R,t}. \quad (11)$$

- Substitute for \hat{R}_t in Euler equation, replace \hat{y}_t by $\mathbf{IE}_{t-1}[\hat{y}_t] + \eta_t^y$ and $\hat{\pi}_t$ by $\mathbf{IE}_{t-1}[\hat{\pi}_t] + \eta_t^\pi$
- and first focus on the evolution of the conditional expectations:

$$\Gamma_0 \underbrace{\begin{bmatrix} \mathbf{IE}_t[\hat{y}_{t+1}] \\ \mathbf{IE}_t[\hat{\pi}_{t+1}] \end{bmatrix}}_{s_t} = \Gamma_1 \underbrace{\begin{bmatrix} \mathbf{IE}_{t-1}[\hat{y}_t] \\ \mathbf{IE}_{t-1}[\hat{\pi}_t] \end{bmatrix}}_{s_{t-1}} + \Psi \epsilon_t + \Pi \underbrace{\begin{bmatrix} \hat{y}_t - \mathbf{IE}_{t-1}[\hat{y}_t] \\ \hat{\pi}_t - \mathbf{IE}_{t-1}[\hat{\pi}_t] \end{bmatrix}}_{\eta_t}, \quad (12)$$

where $\epsilon_t = \epsilon_{R,t}$.

Example

- The system matrices are

$$\Gamma_0 = \begin{bmatrix} 1 & 1/\tau \\ 0 & \beta \end{bmatrix}, \quad \Gamma_1 = \begin{bmatrix} 1 & \psi_1/\tau \\ -\kappa & 1 \end{bmatrix}, \quad \Psi = \begin{bmatrix} 1/\tau \\ 0 \end{bmatrix}, \quad \Pi = \begin{bmatrix} 1 & \psi_1/\tau \\ -\kappa & 1 \end{bmatrix}.$$

Example

- Use method that is similar to Sims (2002) and has been extended to include sunspot equilibria by Lubik and Schorfheide (2003).
- We premultiply the above system by Γ_0 to solve for ξ_t :

$$s_t = \Gamma_1^* s_{t-1} + \Psi^* \epsilon_t + \Pi^* \eta_t. \quad (13)$$

- We proceed with a Jordan decomposition of $\Gamma_1^* = J\Lambda J^{-1}$. Define $w_t = J^{-1}s_t$ and write:

$$w_t = \Lambda w_{t-1} + J^{-1}\Psi^* \epsilon_t + J^{-1}\Pi^* \eta_t. \quad (14)$$

- If $\mathbf{E}_{t-1}[\epsilon_t] = 0$ then a solution of the LRE model is a function

$$\eta_t = \eta_1 \epsilon_t + \eta_2 \zeta_t, \quad (15)$$

where ζ_t is a vector of sunspot shocks such that w_t is stable.

Example

- We partition

$$w_t = \begin{bmatrix} w_{1,t} \\ w_{2,t} \end{bmatrix}, \quad \Lambda = \begin{bmatrix} \Lambda_{11} & 0 \\ 0 & \Lambda_{22} \end{bmatrix}$$

such that the partitions of w_t and Λ conform and Λ_{22} contains all the explosive eigenvalues of Γ_1^* on its diagonal.

- We write the unstable block of the transformed system as

$$\begin{aligned} w_{2,t} &= \Lambda_2 w_{2,t-1} + [J^{-1}\Psi^*]_2 \epsilon_t + [J^{-1}\Pi^*]_2 \eta_t \\ &= \Lambda_2 w_{2,t-1} + \tilde{\Psi} \epsilon_t + \tilde{\Pi} \eta_t \\ &= \Lambda_2 w_{2,t-1} + z_t. \end{aligned} \tag{16}$$

- For w_t to be non-explosive we need to choose η_t such that $z_t = 0$ for all t .

Example

- In principle we can distinguish three cases:
 1. uniqueness: there exists a unique mapping from the structural shocks into the expectation errors that leads to a stable law of motion for w_t ;
 2. indeterminacy: there are multiple mappings from fundamental shocks and sunspot shocks into the expectation errors;
 3. non-existence: no stable rational expectations solution exists.

Example

- If $\psi_1 > 1$ both eigenvalues are unstable. The only stable solution is $\xi_t = 0$, obtained when

$$\Psi^* \epsilon_t + \Pi^* \eta_t = 0. \quad (17)$$

- Thus,

$$\eta_t = -\Pi^{*-1} \Psi^* \epsilon_t. \quad (18)$$

- It can be verified (using somewhat tedious algebra) that the the law of motion for output, inflation, and interest rates is of the form

$$\begin{bmatrix} y_t \\ \pi_t \\ R_t \end{bmatrix} = \frac{1}{\tau + \kappa\psi_1} \begin{bmatrix} -1 \\ -\kappa \\ \tau \end{bmatrix} \epsilon_{R,t} \quad (19)$$

Solving DSGE Models – Slightly More General

- Relax the assumption that $\mathbf{E}_{t-1}[\epsilon_t] = 0$...
- Again, we write the unstable block of the transformed system as

$$\begin{aligned}
 w_{2,t} &= \Lambda_2 w_{2,t-1} + [J^{-1}\Psi^*]_2 \epsilon_t + [J^{-1}\Pi^*]_2 \eta_t \\
 &= \Lambda_2 w_{2,t-1} + \tilde{\Psi} \epsilon_t + \tilde{\Pi} \eta_t \\
 &= \Lambda_2 w_{2,t-1} + z_t.
 \end{aligned} \tag{20}$$

- Thus, we can write

$$w_{2,t} = \Lambda_2^{-1} w_{2,t+1} - \Lambda_2^{-1} z_{t+1} = - \sum_{j=1}^{\infty} (\Lambda_2^{-1})^j z_{t+j}. \tag{21}$$

- Notice that

$$w_{2,t} = \mathbf{E}_t[w_{2,t}] = \mathbf{E}_{t+1}[W_{2,t}] \tag{22}$$

- Hence,

$$w_{2,t} = \sum_{j=1}^{\infty} (\Lambda_2^{-1})^j \mathbf{IE}_t[z_{t+j}] = \sum_{j=1}^{\infty} (\Lambda_2^{-1})^j \mathbf{IE}_{t+1}[z_{t+j}] \quad (23)$$

- Since the η_t 's are rational expectations errors, $\mathbf{IE}_t[\eta_{t+j}] = 0$ for $j > 0$ and

$$w_{2,t} = \sum_{j=1}^{\infty} (\Lambda_2^{-1})^j \tilde{\Psi} \mathbf{IE}_t[\epsilon_{t+j}] = \Lambda_2^{-1} \tilde{\Pi} \eta_{t+1} + \sum_{j=1}^{\infty} (\Lambda_2^{-1})^j \tilde{\Psi} \mathbf{IE}_{t+1}[\epsilon_{t+j}] \quad (24)$$

- Assuming that $\tilde{\Pi}$ is invertible we can solve for η_{t+1}

$$\eta_{t+1} = \tilde{\Pi}^{-1} \sum_{j=1}^{\infty} (\Lambda_2^{-1})^{j-1} \tilde{\Psi} (\mathbf{IE}_t[\epsilon_{t+j}] - \mathbf{IE}_{t+1}[\epsilon_{t+j}]) \quad (25)$$

- It can now be easily verified that $\mathbf{IE}_t[\eta_{t+1}] = 0$. If $\mathbf{IE}_t[\epsilon_{t+j}] = 0$ for $j > 0$ then the condition simplifies to

$$\eta_{t+1} = -\tilde{\Pi}^{-1} \tilde{\Psi} \epsilon_{t+1} \quad (26)$$

which implies that $w_{2,t} = 0$ for all t , provided $w_{2,0} = 0$.

- Notice that once we have an expression for η_t in terms of the structural shock, we can substitute that expression into (7) to obtain a vector autoregressive law of motion for s_t .
- The general case is analyzed in Sims (2002). In particular, the assumption that Γ_0 is invertible is relaxed and the Jordan decomposition is replaced by a General Complex Schur decomposition.

Solving DSGE Models – Nonlinear Methods

- Second-order Perturbation Methods: Judd (1998), Collard and Juillard (2001), Jin and Judd (2002), Schmitt-Grohe and Uribe (2004), Kim, Kim, Schaumburg, Sims (2005), Swanson, Anderson, and Levin (2005).
- Judd's (1998) book covers various global approximation schemes.